# A numerical approach for a system of transport equations in the field of radiotherapy 

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#### Abstract

Numerical schemes for systems of transport equations are commonly constrained by a stability condition of Courant-Friedrichs-Lewy (CFL) type. A system modeling the steady transport of photons and electrons in the field of radiotherapy is considered. Naive discretizations of such a system are commonly constrained by a very restrictive CFL conditions. This issue is circumvented by constructing an implicit scheme based on a relaxation approach. An entropy-based moment model, namely the $M_{1}$ model, is considered. The $M_{1}$ system of equations possesses the non-linear flux terms of an hyperbolic system but no time derivative. The flux terms are well-defined only under a condition on the unknowns, called realizability, which corresponds to the positivity of an underlying kinetic distribution function. The present numerical approach is applicable to non-linear systems which possess no hyperbolic operator, and it preserves the realizability property. However the discrete equations are non-linear and we propose a numerical method to solve such non-linear systems. Our approach is tested on academic and practical cases in 1D, 2D and 3D and it is shown to require significantly less computational power than reference methods.


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## 1 Introduction

The present work aims to construct a numerical solver for systems of steady transport equations emerging in the field of radiotherapy. It is a follow-up to ( $[5,22,52]$ ) and it analyses the numerical methods used $[6,14,48-51,53]$.

[^0]The motion of energetic particles in radiotherapy can be modeled by a system of coupled linear kinetic equations over the fluences of the particles, i.e. over the densities, or distribution functions, of the particles in a phase space composed of position $x \in \mathbb{R}^{3}$, energy $\epsilon \in \mathbb{R}^{+}$and direction of flight $\Omega \in \mathcal{S}^{2}$ on the unit sphere. Due to the high dimensionality of the phase space, solving directly such systems of equations, through either Monte Carlo methods ( $[15,32,35]$ ) or discrete ordinates methods ( [41]; see also [42] and references therein for a review on numerical approaches for dose computation) commonly requires much higher numerical powers than the standard available in medical centers. Recent technological advances lead to the development of industrial codes based on those methods which require considerably lower computational power, i.e. the so-called fast Monte Carlo methods (see e.g. [62]) and Acuros ${ }^{\circledR}$ code ( $[25,46,60]$ ).

As an alternative in this paper, an angular moment extraction technique is used. The resulting system is underdetermined and an entropy minimization procedure is used, leading to the so-called $M_{1}$ model. Such a closure was preferred as it is known to preserve the main features of the underlying kinetic model (especially positivity, hyperbolicity and entropy dissipation), and it models accurately beams of particles. This method was widely used for diverse applications in physics and biology e.g. in astrophysics ( [16, $17,28]$ ), radiative transfer ( [ 20,54$]$ ), in fluid dynamics ( $[27,40,43]$ ), for semi-conductors ( $[29,55]$ ) or chemotaxis ( $[7]$ ) modeling, and showed a considerable reduction of the numerical costs.

Numerical approaches for solving moment equations are typically constrained by a stability condition. Such a condition becomes very restrictive when considering low density media. Typically, the step size (see $[5,52]$ or Section 3 below) needs to be taken proportional to the minimum density in the medium and therefore many steps are required. This problem was first studied for application in radiotherapy in [5] and it was circumvented by the use of a clever change of variables. The previous work [52] showed another approach based on a relaxation method (based on [1,11,45], see also recent work [18]) and on the method of characteristics. However, both those approaches are inappropriate to model the motion of photons. Indeed, those numerical schemes are applicable only to hyperbolic systems, but the transport of photons was shown to be ill-modeled by such equations ( [47]).

In the present paper, an implicit scheme is presented, based on a relaxation method, preserving the realizability property and efficient with large steps. However the discretized equations are non-linear and an iterative solver is constructed to solve such equations.

The paper is constructed as follows. In the next section, models of transport of photons and electrons are presented, first a kinetic model, then the angular moment extraction is described. A first numerical scheme is described for 1D problems in Section 3, an iterative algorithm adapted to this scheme is constructed and tested on an academic test case. This numerical scheme is completed and adapted to multi-D problems in Section 4 and tested on academic test cases in 2D and 3D. Section 5 is devoted to conclusion.

## 2 Models of transport of photons and electrons

Photons and electrons are characterized their position $x \in \mathbb{R}^{3}$, their energy $\epsilon \in \mathbb{R}^{+}$and their direction of flight $\Omega \in \mathcal{S}^{2}$ on the unit sphere. The transported particles are assumed to interact only with the atoms of the background medium. The influence of such interactions on the medium are neglected. In particular, the density of atoms in the medium is given constant data.

### 2.1 A kinetic model

The motion of transported photons $\gamma$ and electrons $e$ is modeled by their fluence $\psi_{\gamma}$ and $\psi_{e}$, i.e. their density in the ( $x, \epsilon, \Omega$ ) space. The fluences satisfy the following steady kinetic equations (see e.g. [24,32])

$$
\begin{align*}
\Omega . \nabla_{x} \psi_{\gamma}(x, \epsilon, \Omega) & =\rho(x)\left[Q_{\gamma \rightarrow \gamma}\left(\psi_{\gamma}\right)(x, \epsilon, \Omega)+Q_{e \rightarrow \gamma}\left(\psi_{e}\right)(x, \epsilon, \Omega)\right],  \tag{2.1a}\\
\Omega . \nabla_{x} \psi_{e}(x, \epsilon, \Omega) & =\rho(x)\left[Q_{e \rightarrow e}\left(\psi_{e}\right)(x, \epsilon, \Omega)+Q_{\gamma \rightarrow e}\left(\psi_{\gamma}\right)(x, \epsilon, \Omega)\right], \tag{2.1b}
\end{align*}
$$

composed of time-independent free transport terms on the left-hand side and collisions operators on the right-hand side. The collision operator $Q_{\alpha \rightarrow \beta}$ models the variations of the fluence $\psi_{\beta}$ due to the collisions involving incident particles $\alpha$. As a first approximation, the influence of the composition of the medium on the collisions is neglected. The collision operator is only chosen to be proportional to the relative density $\rho>0$ compared to the density of water.

The considered collision operators are given by

$$
\begin{align*}
Q_{\gamma \rightarrow \gamma}\left(\psi_{\gamma}\right) & =\left[G_{\gamma \rightarrow \gamma}-P_{\gamma}\right]\left(\psi_{\gamma}\right)  \tag{2.2a}\\
Q_{\gamma \rightarrow e}\left(\psi_{\gamma}\right) & =G_{\gamma \rightarrow e}\left(\psi_{\gamma}\right)  \tag{2.2b}\\
Q_{e \rightarrow \gamma}\left(\psi_{e}\right) & =0,  \tag{2.2c}\\
Q_{e \rightarrow e}\left(\psi_{e}\right) & =\partial_{\epsilon}\left(S \psi_{e}\right)+\left[G_{e \rightarrow e}-P_{e}\right]\left(\psi_{e}\right), \tag{2.2d}
\end{align*}
$$

where the terms $G_{\alpha \rightarrow \beta}$ and $P_{\beta}$ are linear Boltzmann gain and loss terms given by

$$
\begin{align*}
G_{\alpha \rightarrow \beta}\left(\psi_{\alpha}\right)(x, \epsilon, \Omega) & =\int_{\epsilon}^{\epsilon_{\max }} \int_{\mathcal{S}^{2}} \sigma_{\alpha \rightarrow \beta}\left(\epsilon^{\prime}, \epsilon, \Omega^{\prime} . \Omega\right) \psi_{\alpha}\left(x, \epsilon^{\prime}, \Omega^{\prime}\right) d \epsilon^{\prime} d \Omega^{\prime},  \tag{2.2e}\\
P_{\beta}\left(\psi_{\beta}\right)(x, \epsilon, \Omega) & =\sigma_{T, \beta}(\epsilon) \psi_{\beta}(x, \epsilon, \Omega) . \tag{2.2f}
\end{align*}
$$

The stopping power $S>0$, the differential cross sections $\sigma_{\alpha \rightarrow \beta} \geq 0$ and the total cross sections $\sigma_{T, \beta}>0$ are a priori given functions characterizing the collisions in a medium. As we have assumed that the composition of the medium did not influence the collision behaviour, the coefficients $S, \sigma_{T, \beta}$ and $\sigma_{\alpha \rightarrow \beta}$ do not depend on position $x$. Only the density $\rho$ in front of the collision operators in (2.1) depends on $x$. Those collision operators (2.2) were chosen because they accurately model Compton, Møller, and elastic nuclear
scattering (see e.g. [24,32]) which are the predominant effects at the considered ranges of energy.

In medical physics, the function of interest is the energy deposited by the particles per mass unit, so-called dose. At the kinetic level, this dose $D$ is simply given by

$$
\begin{equation*}
D(x)=\int_{\epsilon_{\min }}^{\epsilon_{\max }} \int_{\mathcal{S}^{2}}(-\epsilon) \sum_{\alpha, \beta=\gamma, e} Q_{\alpha \rightarrow \beta}\left(\psi_{\alpha}\right)(x, \epsilon, \Omega) d \Omega d \epsilon . \tag{2.3}
\end{equation*}
$$

### 2.2 A moment model

Solving numerically kinetic equations of the form (2.1-2.2) requires high computational power. Instead, the method of moments is used as it requires lower computational power (see e.g. comparisons in the previous work [49,51,52]).

In the following, the construction of the $M_{1}$ model associated to the kinetic model (2.1-2.2) is recalled.

The moments $\psi^{0}, \psi^{1}$ and $\psi^{2}$ of a fluence $\psi$ are defined by

$$
\begin{gather*}
\psi^{0}(x, \epsilon)=\int_{\mathcal{S}^{2}} \psi(x, \epsilon, \Omega) d \Omega, \quad \psi^{1}(x, \epsilon)=\int_{\mathcal{S}^{2}} \Omega \psi(x, \epsilon, \Omega) d \Omega,  \tag{2.4}\\
\psi^{2}(x, \epsilon)=\int_{\mathcal{S}^{2}} \Omega \otimes \Omega \psi(x, \epsilon, \Omega) d \Omega .
\end{gather*}
$$

Extracting the moments of (2.1) up to order 1 yields, for $i=0,1$,

$$
\begin{align*}
\nabla_{x} \cdot \psi_{\gamma}^{i+1}(x, \epsilon) & =\rho(x)\left[Q_{\gamma \rightarrow \gamma}^{i}\left(\psi_{\gamma}^{i}\right)+Q_{e \rightarrow \gamma}^{i}\left(\psi_{e}^{i}\right)\right](x, \epsilon),  \tag{2.5a}\\
\nabla_{x} \cdot \psi_{e}^{i+1}(x, \epsilon) & =\rho(x)\left[Q_{e \rightarrow e}^{i}\left(\psi_{e}^{i}\right)+Q_{\gamma \rightarrow e}^{i}\left(\psi_{\gamma}^{i}\right)\right](x, \epsilon), \tag{2.5b}
\end{align*}
$$

and the moments of the collision operators of order $i$ are

$$
\begin{align*}
Q_{\gamma \rightarrow \gamma}^{i}\left(\psi_{\gamma}\right) & =\left[G_{\gamma \rightarrow \gamma}^{i}-P_{\gamma}^{i}\right]\left(\psi_{\gamma}^{i}\right),  \tag{2.5c}\\
Q_{\gamma \rightarrow e}^{i}\left(\psi_{\gamma}\right) & =G_{\gamma \rightarrow e}^{i}\left(\psi_{\gamma}^{i}\right),  \tag{2.5d}\\
Q_{e \rightarrow \gamma}^{i}\left(\psi_{e}\right) & =0,  \tag{2.5e}\\
Q_{e \rightarrow e}^{i}\left(\psi_{e}\right) & =\partial_{\epsilon}\left(S \psi_{e}^{i}\right)+\left[G_{e \rightarrow e}^{i}-P_{e}^{i}\right]\left(\psi_{e}^{i}\right), \tag{2.5f}
\end{align*}
$$

where the terms $G_{\alpha \rightarrow \beta}^{i}$ and $P_{\beta}^{i}$ read

$$
\begin{align*}
G_{\alpha \rightarrow \beta}^{i}\left(\psi_{\alpha}^{i}\right)(x, \epsilon) & =\int_{\epsilon}^{\epsilon_{\max }} \sigma_{\alpha \rightarrow \beta}^{i}\left(\epsilon^{\prime}, \epsilon\right) \psi_{\alpha}^{i}\left(x, \epsilon^{\prime}\right) d \epsilon^{\prime}  \tag{2.5~g}\\
P_{\beta}^{i}\left(\psi_{\beta}^{i}\right)(x, \epsilon) & =\sigma_{T, \beta}(\epsilon) \psi_{\beta}^{i}(x, \epsilon)  \tag{2.5h}\\
\sigma_{\alpha \rightarrow \beta}^{i}\left(\epsilon^{\prime}, \epsilon\right) & =2 \pi \int_{-1}^{+1} \mu^{i} \sigma_{\alpha \rightarrow \beta}\left(\epsilon^{\prime}, \epsilon, \mu\right) d \mu \tag{2.5i}
\end{align*}
$$

Remark 2.1. The value of the fluences $\psi_{\gamma}$ and $\psi_{e}$ at all energy $\epsilon^{\prime} \in\left[\epsilon, \epsilon_{\text {max }}\right]$ is required to compute the integrals in the collision operators $(2.5 \mathrm{~g})$ at energy $\epsilon$. In practice, the equations are solved from a maximum energy $\epsilon_{\max }$ to a minimum energy $\epsilon_{\text {min }}$.

The system (2.5) is underdetermined and one more relation, so-called $M_{1}$ closure ( [40, 44]), is added to close the system. This relation expresses $\psi_{\alpha}^{2}$ as a function of $\psi_{\alpha}^{0}$ and $\psi_{\alpha}^{1}$ for $\alpha=\gamma, e$. The $M_{1}$ closure consists in reconstructing the unique function ( $[8-10,30,31,40,57]$ ) of the form

$$
\begin{equation*}
\psi_{M_{1}}(\Omega)=\exp (\lambda \cdot \mathbf{m}(\Omega)), \quad \mathbf{m}(\Omega)=\left(1, \Omega_{1}, \Omega_{2}, \Omega_{3}\right), \quad \lambda=\left(\lambda_{0}, \lambda_{1}, \lambda_{2}, \lambda_{3}\right), \tag{2.6}
\end{equation*}
$$

i.e. by computing the unique vector $\lambda \in \mathbb{R}^{4}$ such that

$$
\begin{equation*}
\int_{\mathcal{S}^{2}} \psi_{M_{1}}(\Omega) d \Omega=\psi^{0}, \quad \int_{\mathcal{S}^{2}} \Omega \psi_{M_{1}}(\Omega) d \Omega=\psi^{1} . \tag{2.7}
\end{equation*}
$$

Then the last moment $\psi^{2}$ is expressed as

$$
\begin{equation*}
\psi^{2}=\int_{\mathcal{S}^{2}} \Omega \otimes \Omega \psi_{M_{1}}(\Omega) d \Omega \tag{2.8}
\end{equation*}
$$

Using (2.6), the closure (2.8) can be rewritten (see e.g. [39,51])

$$
\begin{align*}
\psi^{2} & =\psi^{0}\left(\frac{1-\chi}{2} I d+\frac{3 \chi-1}{2} \frac{\psi^{1} \otimes \psi^{1}}{\left|\psi^{1}\right|^{2}}\right)  \tag{2.9}\\
\chi & =1+\frac{2}{|\alpha|}(1+\operatorname{coth}(|\alpha|)), \quad \alpha=\sqrt{\lambda_{1}^{2}+\lambda_{2}^{2}+\lambda_{3}^{2}}
\end{align*}
$$

where $\chi$ is a scalar called Eddington factor depending on one unique scalar $\alpha$. Using (2.6) and (2.7), one computes

$$
\frac{\left|\psi^{1}\right|}{\psi^{0}}=\frac{|\alpha| \operatorname{coth}(|\alpha|)-1}{|\alpha|} .
$$

which can be inverted to compute $\alpha$ and therefore to obtain the closure (2.9).
Definition 2.1. The $M_{1}$ closure is defined only if there exists a function of the form (2.6) satisfying (2.7). Let us define the following set $\mathcal{R}$ called the realizability domain ( [34])

$$
\begin{equation*}
\left(\psi^{0}, \psi^{1}\right) \in \mathcal{R}=\left\{\left(f^{0}, f^{1}\right) \in \mathbb{R}^{4}, \quad \text { s.t. } \quad\left|f^{1}\right| \leq f^{0}\right\} \cup\left(0,0_{\mathbb{R}^{3}}\right) \tag{2.10}
\end{equation*}
$$

The realizability domain $\mathcal{R}$ is the closure set in $\mathbb{R}^{4}$ of the moments of the functions of the form (2.6). It was shown that the moments of all non-negative integrable functions ( $[8-10]$ ) and measures ( $[19,34]$ ) belonged to the set $\mathcal{R}$.

The requirement (2.10) needs to be kept in mind when constructing numerical schemes because the flux in (2.5) becomes ill-defined if this property is violated. In particular, in the next section, the following remark is used to proof that a numerical scheme preserves the realizability property.
Remark 2.2. The realizability domain is a convex cone. Therefore, any scheme constructed based on a positive combination of realizable vector preserves the realizability property.

Since the differential cross sections $\sigma_{\alpha \rightarrow \beta}$ are also positive scalars, their moments need to satisfy the following realizability condition ( [34])

$$
\begin{equation*}
\left|\sigma_{\alpha \rightarrow \beta}^{1}\right| \leq \sigma_{\alpha \rightarrow \beta}^{0} . \tag{2.11}
\end{equation*}
$$

## 3 A discretization for 1D problems

We aim to construct a solver for the moment system (2.5).
For the sake of simplicity, the numerical approach is described for problems in one spatial dimension. The results are generalized to multidimensional problems in the next section.

In the first subsection, the main problem with the discretization of (2.5) is presented. A first numerical method is described and tested in the remaining subsections.

### 3.1 The fast characteristics problem in 1D

First the problem (2.5) is rewritten, then the main difficulty with the discretization of the rewritten equations is presented.

### 3.1.1 Problem settings

In 1D, the system (2.5) can be rewritten under the vectorial form

$$
\begin{equation*}
\partial_{x} \mathbf{F}(\psi)(x, \epsilon)=\rho(x) \mathbf{Q}(\psi)(x, \epsilon), \tag{3.1a}
\end{equation*}
$$

where the unknown is $\psi=\left(\boldsymbol{\psi}_{\gamma}, \boldsymbol{\psi}_{e}\right) \in \mathcal{R}^{2}$, and $\psi_{\alpha}=\left(\psi_{\alpha}^{0}, \psi_{\alpha}^{1}\right) \in \mathcal{R}$ are the moments of the fluence of particles $\alpha=\gamma, e$. The fluxes $\mathbf{F}(\boldsymbol{\psi})$ and the collision operator $\mathbf{Q}(\boldsymbol{\psi})$ are defined over $\mathcal{R}^{2}$ by

$$
\begin{align*}
& \mathbf{F}(\boldsymbol{\psi})=\left(\mathbf{F}_{M_{1}}\left(\boldsymbol{\psi}_{\gamma}\right), \quad \mathbf{F}_{M_{1}}\left(\boldsymbol{\psi}_{e}\right)\right),  \tag{3.1b}\\
& \mathbf{Q}(\boldsymbol{\psi})=\left(\mathbf{Q}_{\gamma \rightarrow \gamma}\left(\boldsymbol{\psi}_{\gamma}\right)+\mathbf{Q}_{e \rightarrow \gamma}\left(\boldsymbol{\psi}_{e}\right), \quad \mathbf{Q}_{e \rightarrow e}\left(\boldsymbol{\psi}_{e}\right)+\mathbf{Q}_{\gamma \rightarrow e}\left(\boldsymbol{\psi}_{\gamma}\right)\right), \tag{3.1c}
\end{align*}
$$

and are composed of the moments of the kinetic flux and collision operator according to

$$
\begin{align*}
\mathbf{F}_{M_{1}}\left(\boldsymbol{\psi}_{\alpha}\right) & =\left(\psi_{\alpha}^{1} \quad \psi_{\alpha}^{2}\right),  \tag{3.1d}\\
\mathbf{Q}_{\alpha \rightarrow \beta}\left(\psi_{\alpha}\right) & =\left(Q_{\alpha \rightarrow \beta}^{0}\left(\psi_{\alpha}^{0}\right), \quad Q_{\alpha \rightarrow \beta}^{1}\left(\psi_{\alpha}^{1}\right)\right), \tag{3.1e}
\end{align*}
$$

where $\psi_{\alpha}^{2}$ is given by the closure relation (2.8). Before describing the numerical approaches, the following notations are introduced.

Notation 1. The superscript $n$ refers to the discretization in energy $\epsilon$ and the subscript $l$ to the discretization in the $x$ variable. In the next section, the subscript $m$ will refer to the discretization in the second space variable $y$.

According to Remark 2.1, the energy grid is such that $\epsilon^{n}>\epsilon^{n+1}$.

### 3.1.2 Position of the problem

Standard methods to solve (3.1a) create stiff terms at the discrete level, and are therefore very time-consuming for practical applications in medical physics. Such stiffness arises in weakly collisional media, e.g. when the background medium has a low density $\rho$. This problem was illustrated in [5,52] through a 1D electron transport equation of the form

$$
\partial_{x} \mathbf{F}_{M_{1}}\left(\boldsymbol{\psi}_{e}\right)=\rho\left[\partial_{\epsilon}\left(S \boldsymbol{\psi}_{e}\right)+A \boldsymbol{\psi}_{e}\right], \quad A=\left(\begin{array}{cc}
0 & 0  \tag{3.2}\\
0 & T
\end{array}\right),
$$

with $T \in \mathbb{R}^{+}$by using the scheme

$$
\begin{equation*}
\frac{\mathbf{F}_{e, l+\frac{1}{2}}^{n}-\mathbf{F}_{e, l-\frac{1}{2}}^{n}}{\Delta x}-\rho_{l} \frac{S^{n} \boldsymbol{\psi}_{e, l}^{n}-S^{n+1} \boldsymbol{\psi}_{e, l}^{n+1}}{\Delta \epsilon^{n}}=\rho_{l} A^{n} \boldsymbol{\psi}_{e, l}^{n} \tag{3.3a}
\end{equation*}
$$

with numerical fluxes of Lax-Friedrichs type ( $[38,58]$ )

$$
\begin{equation*}
\mathbf{F}_{e, l+\frac{1}{2}}^{n}=\frac{1}{2}\left[\mathbf{F}_{M_{1}}\left(\psi_{e, l+1}^{n}\right)+\mathbf{F}_{M_{1}}\left(\boldsymbol{\psi}_{e, l}^{n}\right)+\left(\boldsymbol{\psi}_{e, l+1}^{n}-\boldsymbol{\psi}_{e, l}^{n}\right)\right] . \tag{3.3b}
\end{equation*}
$$

Such a scheme was shown to be consistent with (3.2), however it is only stable under the following Courant-Friedrichs-Lewy (CFL) condition

$$
\begin{equation*}
\Delta \epsilon^{n} \leq S^{n} \Delta x \min _{l}\left(\rho_{l}\right) . \tag{3.4}
\end{equation*}
$$

The condition (3.4) turns very restrictive when considering low collisional media, here when $\rho$ is small. In such a case, one requires a very large number of energy steps and therefore considerably long computational times are necessary.

A first solution to this problem was proposed in [5] by the use of a change of variable. An alternative was proposed in [52] through a method of characteristic applied on a relaxed system for (3.2). However, both methods can only be used on hyperbolic systems, and are therefore not applicable to the coupled photons-electrons transport equations (2.5).

The present discretization of the system (3.1) is presented in three parts.

- Step 1 (Subsection 3.2): the advection operator is discretized, this corresponds to the discretization over the position variable $x$.
- Step 2 (Subsection 3.3): the collision operator is discretized, this corresponds to the discretization over the energy variable $\epsilon$.
- Step 3 (Subsection 3.4): both discretizations are gathered to construct a numerical scheme and an iterative algorithm is constructed to solve the resulting discrete equations.


### 3.2 Discretization of the advection operator

In the spirit of [52], a numerical scheme for (2.5) is constructed based on the relaxation method developed in $[1,11,12,45]$. The principle of the relaxation method is first recalled.

The relaxation method is presented to justify the construction of the implicit scheme (3.11) below. The aim is to isolate the non-linearity in (3.1a) into a relaxation term and to solve easierly the remaining linear part by the use of an implicit scheme.

A vectorial BGK approximation of the system (2.5) is performed. In practice, the nonlinear flux term is replaced by a linear advection term and a relaxation term is added the right-hand side. Let us introduce the following system of two relaxed equations

$$
\begin{align*}
c^{-} \partial_{x} \mathbf{f}_{\tau}^{-}-\rho \mathbf{Q}\left(\mathbf{f}_{\tau}^{-}\right) & =\frac{\mathbf{M}^{-}-\mathbf{f}_{\tau}^{-}}{\tau},  \tag{3.5a}\\
c^{+} \partial_{x} \mathbf{f}_{\tau}^{+}-\rho \mathbf{Q}\left(\mathbf{f}_{\tau}^{+}\right) & =\frac{\mathbf{M}^{+}-\mathbf{f}_{\tau}^{+}}{\tau}, \tag{3.5b}
\end{align*}
$$

where $\mathbf{f}_{\tau}^{ \pm}$are the unknowns relaxing toward the equilibrium represented by Maxwellians $\mathbf{M}^{ \pm}(\boldsymbol{\psi}) \in \mathcal{R}^{2}$, and $\tau$ is a relaxation parameter. The unknowns $\mathbf{f}_{\tau}^{ \pm}$are only related to the original $\psi$ in the limit $\tau \rightarrow 0$.

The Maxwellians $\mathbf{M}^{ \pm}(\boldsymbol{\psi}) \in \mathcal{R}^{2}$ are chosen to relate to the original system through the following consistency formulae

$$
\begin{equation*}
\mathbf{M}^{+}+\mathbf{M}^{-}=\boldsymbol{\psi} \in \mathcal{R}^{2}, \quad c^{+} \mathbf{M}^{+}+c^{-} \mathbf{M}^{-}=\mathbf{F}(\boldsymbol{\psi}) . \tag{3.6a}
\end{equation*}
$$

The relaxation velocities $c^{ \pm}(\psi) \in \mathbb{R}$ are chosen such that they bound the physical velocities. This leads to the following stability requirement ( $[1,11,12]$ )

$$
\begin{equation*}
\operatorname{Sp}\left(\mathbf{F}^{\prime}(\boldsymbol{\psi})\right) \subset\left[c^{-}, c^{+}\right] . \tag{3.7}
\end{equation*}
$$

Formally, at the limit $\tau \rightarrow 0$ in (3.5), one obtains $\mathbf{f}_{0}^{ \pm}=\mathbf{M}^{ \pm}$. Then replacing $\mathbf{f}^{ \pm}$by $\mathbf{M}^{ \pm}$ and summing the two equations (3.5) yields (3.1). Therefore, one recovers the solution of the original equation (3.1) in the limit case $\tau \rightarrow 0$ as

$$
\begin{equation*}
\boldsymbol{\psi}=\lim _{\tau \rightarrow 0}\left(\mathbf{f}_{\tau}^{+}+\mathbf{f}_{\tau}^{-}\right) \tag{3.8}
\end{equation*}
$$

We refer to $[1,2,11,12,45]$ for a proper analysis of this asymptotic limit.
For the sake of simplicity, in the 1D case, we use the following classical result (see e.g. $[4,13,52])$, to choose the relaxation parameters.

Lemma 3.1. The eigenvalues of the Jacobian $\mathbf{F}_{M_{1}}^{\prime}\left(\boldsymbol{\psi}_{\alpha}\right)$ of the $M_{1}$ fluxes are bounded by 1 for all $\boldsymbol{\psi}_{\alpha} \in \mathcal{R}$, that is

$$
\left.\forall \boldsymbol{\psi}_{\alpha} \in \mathcal{R}, \quad \operatorname{Sp}\left(\mathbf{F}_{M_{1}}^{\prime}\left(\boldsymbol{\psi}_{\alpha}\right)\right) \subset\right]-1,1[.
$$

Furthermore, for all realizable moments $\psi_{\alpha} \in \mathcal{R}$, one has

$$
\boldsymbol{\psi}_{\alpha} \pm \mathbf{F}_{M_{1}}\left(\boldsymbol{\psi}_{\alpha}\right) \in \mathcal{R}
$$

Thus, in 1D, we will use the following parameters

$$
\begin{align*}
c^{+}= & =-c^{-}, \quad \mathbf{M}^{ \pm}=\frac{\boldsymbol{\psi} \pm \mathbf{F}(\boldsymbol{\psi})}{2} \in \mathcal{R}^{2}  \tag{3.9a}\\
& \pm \partial_{x} \mathbf{f}_{\tau}^{ \pm}-\rho \mathbf{Q}\left(\mathbf{f}_{\tau}^{ \pm}\right)=\frac{\mathbf{M}^{ \pm}-\mathbf{f}_{\tau}^{ \pm}}{\tau} \tag{3.9b}
\end{align*}
$$

which satisfy the requirements (3.6).
Then we use upwind fluxes on (3.9b) leading to the scheme

$$
\begin{equation*}
\frac{\mathbf{f}_{\tau, l}^{ \pm, n}-\mathbf{f}_{\tau, l \mp 1}^{ \pm, n}}{\Delta x}-\rho \mathbf{Q}\left(\mathbf{f}_{\tau, l}^{ \pm, n}\right)=\frac{\mathbf{M}^{ \pm}-\mathbf{f}_{\tau, l}^{ \pm, n}}{\tau} \tag{3.10}
\end{equation*}
$$

Then summing these equations over $\pm$ and having $\tau \rightarrow 0$ leads to define the following scheme over $\boldsymbol{\psi}$

$$
\begin{gather*}
\frac{\mathbf{F}_{l+\frac{1}{2}}^{n}-\mathbf{F}_{l-\frac{1}{2}}^{n}}{\Delta x}-[\rho \mathbf{Q}(\boldsymbol{\psi})]_{l}^{n}=0  \tag{3.11a}\\
\mathbf{F}_{l+\frac{1}{2}}^{n}=  \tag{3.11b}\\
\frac{1}{2}\left[\mathbf{F}\left(\boldsymbol{\psi}_{l+1}^{n}\right)+\mathbf{F}\left(\boldsymbol{\psi}_{l}^{n}\right)-\left(\boldsymbol{\psi}_{l+1}^{n}-\boldsymbol{\psi}_{l}^{n}\right)\right] .
\end{gather*}
$$

The term $[\rho \mathbf{Q}(\boldsymbol{\psi})]_{l}^{n}$ will be defined in the next subsection.
Remark 3.1. Since the collision operator $\mathbf{Q}$ is linear, one should recover

$$
\mathbf{Q}(\boldsymbol{\psi})_{l}^{n}=\mathbf{Q}\left(\mathbf{M}^{-}+\mathbf{M}^{+}\right)_{l}^{n}=\mathbf{Q}\left(\mathbf{M}^{-}\right)_{l}^{n}+\mathbf{Q}\left(\mathbf{M}^{+}\right)_{l}^{n}
$$

### 3.3 Discretization of the collision operator

The collision terms are simply discretized with a quadrature rule for the integrals in $\epsilon$ and an implicit Euler discretization for the term $\partial_{\epsilon}\left(S \boldsymbol{\psi}_{e}\right)$. This reads

$$
\begin{align*}
{[\rho \mathbf{Q}(\boldsymbol{\psi})]_{l}^{n} } & =\rho_{l} \mathbf{Q}(\boldsymbol{\psi})_{l}^{n},  \tag{3.12a}\\
\mathbf{Q}(\boldsymbol{\psi})_{l}^{n} & =\left(\mathbf{Q}_{\gamma \rightarrow \gamma}\left(\boldsymbol{\psi}_{\gamma}\right)_{l}^{n}+\mathbf{Q}_{e \rightarrow \gamma}\left(\boldsymbol{\psi}_{e}\right)_{l}^{n}, \quad \mathbf{Q}_{e \rightarrow e}\left(\boldsymbol{\psi}_{e}\right)_{l}^{n}+\mathbf{Q}_{\gamma \rightarrow \gamma}\left(\boldsymbol{\psi}_{\gamma}\right)_{l}^{n}\right),  \tag{3.12b}\\
\mathbf{Q}_{\alpha \rightarrow \beta}\left(\boldsymbol{\psi}_{\alpha}\right)_{l}^{n} & =\left(Q_{\alpha \rightarrow \beta}^{0}\left(\psi_{\alpha}^{0}\right)_{l}^{n}, \quad Q_{\alpha \rightarrow \beta}^{1}\left(\psi_{\alpha}^{1}\right)_{l}^{n}\right), \tag{3.12c}
\end{align*}
$$

where each discrete collision operators is, for $i=0,1$

$$
\begin{align*}
Q_{\gamma \rightarrow \gamma}^{i}\left(\psi_{\gamma}^{i}\right)_{l}^{n} & =\sum_{n^{\prime}=1}^{n} \sigma_{\gamma \rightarrow \gamma}^{i, n^{\prime}, n} \psi_{\gamma, l}^{i, n^{\prime}} \Delta \epsilon^{n^{\prime}}-\sigma_{T, \gamma}^{n} \psi_{\gamma, l^{\prime}}^{i, n}  \tag{3.12d}\\
Q_{\gamma \rightarrow e}^{i}\left(\psi_{\gamma}^{i}\right)_{l}^{n} & =\sum_{n^{\prime}=1}^{n} \sigma_{\gamma \rightarrow e}^{i, n^{\prime}, n} \psi_{\gamma, l}^{i, n^{\prime}} \Delta \epsilon^{n^{\prime}},  \tag{3.12e}\\
Q_{e \rightarrow \gamma}^{i}\left(\psi_{e}^{i}\right)_{l}^{n} & =0,  \tag{3.12f}\\
Q_{e \rightarrow e}^{i}\left(\psi_{e}^{i}\right)_{l}^{n} & =\frac{S^{n-1} \psi_{e, l}^{i, n-1}-S^{n} \psi_{e, l}^{i, n}}{\Delta \epsilon^{n}}+\sum_{n^{\prime}=1}^{n} \sigma_{e \rightarrow e}^{i, n^{\prime}, n} \psi_{e, l}^{i, n^{\prime}} \Delta \epsilon^{n^{\prime}}-\sigma_{T, e}^{n} \psi_{e, l}^{i, n} . \tag{3.12g}
\end{align*}
$$

Remark 3.2. - We choose this particular discretization because it leads to an implicit scheme of the form (3.11-3.12), in the sense that the fluxes $\mathbf{F}_{l+\frac{1}{2}}^{n}$ are evaluated at the latest energy step $\epsilon^{n}$. In practice, the obtained scheme is efficient even without imposing a restriction on the step size $\Delta \epsilon^{n}$, which circumvents the problem presented in Subsection 3.1.

- The discretization is constructed to be of order one in $\Delta x$ and in $\Delta \epsilon^{n}$ when the solution is smooth. So the scheme (3.11-3.12) is consistent with the continuous equation (3.1a).
- In order to use the present scheme, one needs to compute $\psi_{l}^{n}$ for all $l$ based on $\psi_{l}^{n^{\prime}}$ for $n^{\prime}<n$. Here, this implies solving the non-linear equation (3.11-3.12) over the vector $\left(\boldsymbol{\psi}_{l}^{n}\right)_{l=1, \ldots, l_{\text {max }}} \in\left(\mathcal{R}^{2}\right)^{l_{\text {max }}}$.


### 3.4 An iterative solver for the 1D scheme

Writing together the discretization of the 1D advection and the collision term with the relaxation parameters (3.9) yields the following numerical scheme

$$
\begin{equation*}
-\mathbf{L}\left(\boldsymbol{\psi}_{l-1}^{n}\right)+\mathbf{D}\left(\boldsymbol{\psi}_{l}^{n}\right)-\mathbf{U}\left(\boldsymbol{\psi}_{l+1}^{n}\right)=\rho_{l} \mathbf{R}_{l}^{n} \tag{3.13a}
\end{equation*}
$$

where the operators $\mathbf{L}$ and $\mathbf{U}$ are non-linear and $\mathbf{D}$ is linear and invertible. Those operators are given by

$$
\begin{gather*}
\mathbf{L}\left(\boldsymbol{\psi}_{l-1}^{n}\right)=\frac{\boldsymbol{\psi}_{l-1}^{n}+\mathbf{F}\left(\boldsymbol{\psi}_{l-1}^{n}\right)}{2 \Delta x}, \quad \mathbf{U}\left(\boldsymbol{\psi}_{l+1}^{n}\right)=\frac{\boldsymbol{\psi}_{l+1}^{n}-\mathbf{F}\left(\boldsymbol{\psi}_{l+1}^{n}\right)}{2 \Delta x},  \tag{3.13b}\\
\mathbf{D}\left(\boldsymbol{\psi}_{l}^{n}\right)=\left(\frac{I d}{\Delta x}+\rho_{l} A^{n}\right) \boldsymbol{\psi}_{l}^{n} \tag{3.13c}
\end{gather*}
$$

and the matrix $A^{n}$ and the source $\mathbf{R}_{l}^{n}$ are given by

$$
\begin{array}{rlll}
A^{n} \boldsymbol{\psi}_{l}^{n} & =\left(B_{0}^{n} \boldsymbol{\psi}_{l}^{n}, \quad B_{1}^{n} \boldsymbol{\psi}_{l}^{n},\right. & \left.D_{0}^{n} \boldsymbol{\psi}_{l}^{n}, \quad D_{1}^{n} \boldsymbol{\psi}_{l}^{n}\right) \\
\mathbf{R}_{l}^{n} & =\left(\begin{array}{lll}
C_{0}^{n}, & C_{1}^{n}, & E_{0}^{n},
\end{array} E_{1}^{n}\right)+B C_{l}^{n} \tag{3.13e}
\end{array}
$$

with, for $i=0,1$,

$$
\begin{align*}
B_{i}^{n} \boldsymbol{\psi}_{l}^{n} & =\left(\sigma_{T, \gamma}^{n}-\sigma_{\gamma \rightarrow \gamma}^{i, n, n} \Delta \epsilon^{n}\right) \psi_{\gamma, l^{\prime}}^{i, n} \quad C_{i}^{n}=\sum_{n^{\prime}=1}^{n-1} \sigma_{\gamma \rightarrow \gamma}^{i, n^{\prime}, n} \psi_{\gamma, l}^{i, n^{\prime}} \Delta \epsilon^{n^{\prime}}  \tag{3.13f}\\
D_{i}^{n} \boldsymbol{\psi}_{l}^{n} & =\left(\frac{S^{n}}{\Delta \epsilon^{n}}+\sigma_{T, e}^{n}-\sigma_{e \rightarrow e}^{i, n, n} \Delta \epsilon^{n}\right) \psi_{e, l}^{i, n}-\sigma_{\gamma \rightarrow e}^{i, n, n} \psi_{\gamma, l}^{i, n} \Delta \epsilon^{n}  \tag{3.13~g}\\
E_{i}^{n} & =\frac{S^{n-1}}{\Delta \epsilon^{n}} \psi_{e, l}^{i, n-1}+\sum_{n^{\prime}=1}^{n-1}\left(\sigma_{\gamma \rightarrow \gamma}^{i, n^{\prime}, n} \psi_{\gamma, l}^{i, n^{\prime}}+\sigma_{e \rightarrow e}^{i, n^{\prime}, n} \psi_{e, l}^{i, n^{\prime}}\right) \Delta \epsilon^{n^{\prime}} \tag{3.13h}
\end{align*}
$$

Defining properly boundary conditions for moment models based on the underlying kinetic ones remains an open problem (see e.g. [26,36,56] for linear moment equations). For the sake of simplicity, we use here discrete boundary conditions defined as a source term in (3.13e) with

$$
B C_{l}^{n}=\boldsymbol{\psi}_{0}^{n} \delta_{1, l}+\boldsymbol{\psi}_{l_{\max }+1}^{n} \delta_{l_{\max }, l}
$$

with given $\psi_{0}^{n} \in \mathcal{R}^{2}$ and $\psi_{l_{\max }+1}^{n} \in \mathcal{R}^{2}$.
In order to use this scheme, one needs to solve (3.13) at each energy step $n$, which is a non-linear equation on the vector $\left(\psi^{n}\right)_{l=1, \ldots, l_{\max }}$. For this purpose, we propose an iterative solver inspired of [21], which was tested in [6,48,50,51,53].

Algorithm 1. Initialization: At energy step $n$, set $\boldsymbol{\psi}_{l}^{n,(0)}=\boldsymbol{\psi}_{l}^{n-1}$ for all $l$. Iteration: Compute iteratively

$$
\begin{equation*}
\boldsymbol{\psi}_{l}^{n,(k+1)}=\mathbf{D}^{-1}\left(\mathbf{L}\left(\boldsymbol{\psi}_{l-1}^{n,(k)}\right)+\mathbf{U}\left(\boldsymbol{\psi}_{l+1}^{n,(k)}\right)+\rho_{l} \mathbf{R}_{l}^{n}\right) \tag{3.14}
\end{equation*}
$$

until convergence.
Proposition 3.1. Suppose that $\mathbf{R}_{l}^{n} \in \mathcal{R}^{2}$ is realizable for all $l$, and that

$$
\begin{equation*}
\min _{l} \rho_{l} \Delta x m_{A}^{n}>\sup _{\psi \in \mathcal{R}^{2}}\left(\frac{\left(M_{F}-m_{F}\right)(\boldsymbol{\psi})}{2}\right) \tag{3.15}
\end{equation*}
$$

where

$$
m_{A}^{n}=\min S p\left(A^{n}\right), \quad m_{F}(\boldsymbol{\psi})=\min S p\left(\mathbf{F}^{\prime}(\boldsymbol{\psi})\right), \quad M_{F}(\boldsymbol{\psi})=\max S p\left(\mathbf{F}^{\prime}(\boldsymbol{\psi})\right)
$$

are respectively the minimum eigenvalue of the matrix $A^{n}$, of the Jacobian $\mathbf{F}^{\prime}(\boldsymbol{\psi})$ and the maximum eigenvalue of $\mathbf{F}^{\prime}(\boldsymbol{\psi})$.
Then there exists a unique solution $\left(\boldsymbol{\psi}^{n}\right)_{l=1, \ldots, l_{\max }} \in\left(\mathcal{R}^{2}\right)^{l_{\max }}$ satisfying (3.13) for all $l$.
Moreover, Algorithm 1 converges to this solution.
Proof. Define the operator $J$ over $\psi \in\left(\mathcal{R}^{2}\right)^{l_{\text {max }}}$ by

$$
\boldsymbol{\psi}^{n,(k+1)}=J\left(\boldsymbol{\psi}^{n,(k)}\right)
$$

where the $l$-th component $J\left(\psi^{n,(k+1)}\right)_{l}$ is given by (3.14), that is

$$
\begin{equation*}
J\left(\boldsymbol{\psi}^{n_{,}(k+1)}\right)_{l}=\left(\frac{I d}{\Delta x}+\rho_{l} A^{n}\right)^{-1}\left[\rho_{l} \mathbf{R}_{l}^{n}+\frac{\boldsymbol{\psi}_{l+1}^{n_{1}(k)}-\mathbf{F}\left(\boldsymbol{\psi}_{l+1}^{n_{,}(k)}\right)}{2 \Delta x}+\frac{\boldsymbol{\psi}_{l-1}^{n_{l}(k)}+\mathbf{F}\left(\boldsymbol{\psi}_{l-1}^{n_{,}(k)}\right)}{2 \Delta x}\right] . \tag{3.16}
\end{equation*}
$$

First, we verify that $J$ preserves the realizability from one step to another.
Let us suppose $\psi^{n,(k)} \in\left(\mathcal{R}^{2}\right)^{l_{\text {max }}}$. Then $\boldsymbol{\psi}_{l+1}^{n^{\prime,(k)}}-\mathbf{F}\left(\boldsymbol{\psi}_{l+1}^{n_{l},(k)}\right) \in \mathcal{R}^{2}$ and $\boldsymbol{\psi}_{l-1}^{n^{\prime}(k)}+\mathbf{F}\left(\boldsymbol{\psi}_{l-1}^{n_{l},(k)}\right) \in \mathcal{R}^{2}$ are realizable according to the second part of Lemma 3.1. Thus the term between square brackets in (3.16) is realizable according to Remark 2.2.

Now, we need to prove that the operator $\left(\frac{I d}{\Delta x}+\rho_{l} A^{n}\right)^{-1}$ preserves the realizability property. Using its definition (3.13), the matrix $A^{n}$ can be rewritten

$$
A^{n}=\left(\begin{array}{cccc}
a_{0} & 0 & 0 & 0 \\
0 & a_{1} & 0 & 0 \\
b_{0} & 0 & c_{0} & 0 \\
0 & b_{1} & 0 & c_{1}
\end{array}\right), \quad\left(\frac{I d}{\Delta x}+\rho_{l} A^{n}\right)^{-1}=\left(\begin{array}{cccc}
\alpha_{0} & 0 & 0 & 0 \\
0 & \alpha_{1} & 0 & 0 \\
\beta_{0} & 0 & \gamma_{0} & 0 \\
0 & \beta_{1} & 0 & \gamma_{1}
\end{array}\right)
$$

where

$$
\begin{array}{ll}
a_{i}=\sigma_{T, \gamma}^{n}-\sigma_{\gamma \rightarrow \gamma}^{i, n, n} \Delta \epsilon^{n}, & \alpha_{i}=\frac{1}{\frac{1}{\Delta x}+\rho_{l} a_{i}}, \\
b_{i}=-\sigma_{\gamma \rightarrow e}^{i, n, n} \Delta \epsilon^{n}, & \beta_{i}=-\frac{b_{i}}{\left(\frac{1}{\Delta x}+\rho_{l} a_{i}\right)\left(\frac{1}{\Delta x}+\rho_{l} c_{i}\right)}, \\
c_{i}=\frac{S^{n}}{\Delta \epsilon^{n}}+\sigma_{T, e}^{n}-\sigma_{e}^{i, n, n} \Delta \epsilon^{n}, & \gamma_{i}=\frac{1}{\frac{1}{\Delta x}+\rho_{l} c_{i}} .
\end{array}
$$

Using (2.11) leads to

$$
0 \leq a_{0} \leq a_{1}, \quad 0 \leq c_{0} \leq c_{1} \quad \text { and } \quad-b_{0} \geq-b_{1} \geq 0
$$

and so

$$
\alpha_{0} \geq \alpha_{1}>0, \quad \beta_{0} \geq \beta_{1}>0 \quad \gamma_{0} \geq \gamma_{1}>0 .
$$

Then one verifies that if $\psi \in \mathcal{R}^{2}$, then $\left(\frac{I d}{\Delta x}+\rho_{l} A^{n}\right)^{-1} \psi$ is a positive combination of vectors satisfying the criteria (2.10). So the operator $\left(\frac{I d}{\Delta x}+\rho_{l} A^{n}\right)^{-1}$ preserves the realizability, and $J$ is an operator from $\left(\mathcal{R}^{2}\right)^{l_{\text {max }}}$ into itself.

Now, in order to prove that Algorithm 1 converges, we prove that $J$ is a contraction.

Derivating $J(\boldsymbol{\psi})$ according to $\psi$ reads

$$
\begin{align*}
d_{\psi} J(\boldsymbol{\psi}) & =\left(\begin{array}{ccccc}
0 & J_{1,2}^{n} & 0 & \cdots & 0 \\
J_{2,1}^{n} & 0 & J_{2,3}^{n} & 0 & \vdots \\
0 & \ddots & \ddots & \ddots & 0 \\
\vdots & \ddots & \ddots & \ddots & J_{l_{\max }-1, l_{\max }} \\
0 & \cdots & 0 & J_{l_{\max }, l_{\max }-1}^{0}
\end{array}\right)  \tag{3.17a}\\
J_{l, l-1}^{n} & =\left(\frac{I d}{\Delta x}+\rho_{l} A^{n}\right)^{-1}\left(\frac{I d+\mathbf{F}^{\prime}\left(\boldsymbol{\psi}_{l-1}\right)}{2 \Delta x}\right)  \tag{3.17b}\\
J_{l, l+1}^{n} & =\left(\frac{I d}{\Delta x}+\rho_{l} A^{n}\right)^{-1}\left(\frac{I d-\mathbf{F}^{\prime}\left(\boldsymbol{\psi}_{l+1}\right)}{2 \Delta x}\right) \tag{3.17c}
\end{align*}
$$

so $d_{\psi} J(\boldsymbol{\psi})$ is a block matrix with non-zero blocks on the super- and sub-diagonal.
Using a Gershgorin theorem for block matrices (see e.g. [59]) provides

$$
S p\left(d_{\psi} J(\boldsymbol{\psi})\right) \subset[-r, r],
$$

with a spectral radius satisfying

$$
r \leq \max _{l} \frac{\| \| I d-\mathbf{F}^{\prime}\left(\boldsymbol{\psi}_{l}\right)\| \|+\mid\left\|I d+\mathbf{F}^{\prime}\left(\boldsymbol{\psi}_{l}\right)\right\| \|}{2 \Delta x \min S p\left(\frac{I d}{\Delta x}+\rho_{l} A^{n}\right)} .
$$

Using Lemma 3.1 and the fact that $A^{n}$ is positive definite, one finds upper and lower bounds on the eigenvalues of $I d \pm \mathbf{F}^{\prime}(\boldsymbol{\psi})$ and of $I d+\rho_{l} A^{n} \Delta x$ to find

$$
\begin{equation*}
r \leq \max _{l} \frac{1+\frac{\left(M_{F}-m_{f}\right)\left(\psi_{l}\right)}{2}}{1+\rho_{l} m_{A}^{n} \Delta x} \tag{3.18}
\end{equation*}
$$

thus $r<1$ under condition (3.15), so $J$ is a contraction and Algorithm 1 converges to the unique fixed point of $J$.

Remark 3.3. - The requirement (3.15) corresponds to a CFL-like condition. This condition was added here for to prove that the algorithm converges. However, the bound (3.18) used in the proof is not optimal. We have not yet found any theoretical nor experimental test case violating (3.15) that lead to a non-converging sequence $\left(\boldsymbol{\psi}^{n,(k)}\right)_{k=1, \ldots, \infty}$, even with very low collisional media (for small $\left.\rho_{l}\right)$. In the test cases below, $\Delta x$ and $\Delta \epsilon^{n}$ do not necessarily respect the condition (3.15) and we always verify experimentally the convergence of Algorithm 1 at every step $n$. We refer e.g. to $[3,23,33,61]$ and references therein for more complete study on convergence for such algorithms.

- The iterative method proposed in Algorithm 1 can be interpreted as a Jacobi method with non-linear extradiagonal term. Similarily, Gauss-Seidel and successive overrelaxation (SOR) methods for the non-linear problem (3.13) can also be implemented. For example, the non-linear Gauss-Seidel type method consists in solving alternatively

$$
\begin{align*}
& \boldsymbol{\psi}_{l}^{n,(k+1)}=\mathbf{D}^{-1}\left(\mathbf{L}\left(\boldsymbol{\psi}_{l-1}^{n_{l}(k+1)}\right)+\mathbf{U}\left(\boldsymbol{\psi}_{l+1}^{n_{l}(k)}\right)+\rho_{l} \mathbf{R}_{l}^{n}\right),  \tag{3.19a}\\
& \boldsymbol{\psi}_{l}^{n_{l}(k+1)}=\mathbf{D}^{-1}\left(\mathbf{L}\left(\boldsymbol{\psi}_{l-1}^{n_{1}(k)}\right)+\mathbf{U}\left(\boldsymbol{\psi}_{l+1}^{n_{1}(k+1)}\right)+\rho_{l} \mathbf{R}_{l}^{n}\right), \tag{3.19b}
\end{align*}
$$

in Algorithm 1 instead of (3.14). One can easily adapt the proof of Proposition 3.1 for such algorithms.

- The convergence rate of Algorithm 1 depends on the eigenvalues of $d_{\psi} J(\boldsymbol{\psi})$. In the computations (3.18), the worst possible convergence rate corresponds to the case where $\left(M_{F}-m_{F}\right)\left(\psi_{l}\right)$ has the highest value. Such highest value is obtained in the limit case of a purely anisotropic distribution (see e.g. computations in [4]) modeled by a fluence

$$
\psi(\Omega)=K \delta\left(\Omega_{1}-1\right)
$$

Thus Algorithm 1 is slower if the expected solution of (2.1) possesses purely anisotropic region.

- Algorithm 1 is based on a fixed point theorem. One may think of using Newton's method to accelerate the convergence. Such method is unavailable here since the Jacobian $\mathbf{F}^{\prime}(\boldsymbol{\psi})$ is not computed here. Only its upper bound was used in the proof of Proposition 3.1.


### 3.5 Numerical experiments

This subsection is devoted to study experimentally the convergence of the present method. Especially, two convergence rates are studied:

- The convergence according to the number $k_{\max }$ of iterations in Algorithm 1.
- The convergence according to the $\Delta \epsilon^{n}$ and $\Delta x$ of the numerical scheme (3.13).

Those two convergence rates are studied through the following test case.
We consider a 1D domain $Z=[0 \mathrm{~cm}, 8 \mathrm{~cm}]$ uniformly composed of water (i.e. $\rho=1$ ). A source of electrons is imposed as initial condition at $\epsilon_{0}=5 \mathrm{MeV}$ modeled by the initial condition

$$
\begin{equation*}
\psi_{e}(x, 5 \mathrm{MeV}, \Omega)=K \mathbf{1}_{[3 \mathrm{~cm}, 5 \mathrm{~cm}]}(x), \quad \psi_{\gamma}(x, 5 \mathrm{MeV}, \Omega)=0 \tag{3.20a}
\end{equation*}
$$

and a zero condition flux is considered on the boundary which corresponds to extracting the moments of distributions of the form

$$
\begin{array}{lll}
\text { for } & \Omega_{1}>0, & \psi_{e}(0 \mathrm{~cm}, \epsilon, \Omega)=0,
\end{array} \psi_{\gamma}(0 \mathrm{~cm}, \epsilon, \Omega)=0, ~ 子, ~ \psi_{e}(8 \mathrm{~cm}, \epsilon, \Omega)=0, \quad \psi_{\gamma}(8 \mathrm{~cm}, \epsilon, \Omega)=0 .
$$

Remark that the equations of system (2.1) are decoupled. As we impose no source of photons, and due to the considered physics, no photons are created in the system. Thus the solution of (2.1a) is simply $\psi_{\gamma}=0$, and the discretization (3.13) can be simplified when considering beams of electrons only. This is no longer true when considering more complex physics, e.g. when taking into account Bremsstrahlung effect [37,42].

The mesh is composed of 800 cells in $x$ uniformly distributed. The step size $\Delta \epsilon^{n}$ and the grid in $\epsilon$ are chosen such that

$$
\begin{equation*}
\Delta \epsilon^{n}=5 S^{n} \Delta x . \tag{3.22}
\end{equation*}
$$

This corresponds approximately to fixing

$$
m_{A}^{n} \Delta x=5 .
$$

With such a gridsize, we aim to avoid to have numerical diffusion effects depending $\Delta \epsilon^{n}$.
This test case is academic and aims only to study experimentally the numerical convergence rates. The converged dose normalized by its maximum value computed with Algorithm 1 is provided on Fig. 1 as an indication. For more examples with different applications in medical physics, we refer e.g. to $[6,48,51,53]$.

### 3.5.1 Convergence results of the iterative algorithm

The iterative method of Algorithm 1 requires a criterium to stop.
A first naive criterium consists in fixing the number of iterations $k_{\max }$. This is not optimal, neither in terms of precision nor in terms of computational costs. At each energy step, the desired solution follows (3.13) for all $l$. Then one better stopping criterium consists in defining the residual

$$
\begin{equation*}
r^{n,(k)}=\max _{l}\left\|-\mathbf{L}\left(\boldsymbol{\psi}_{l-1}^{n_{1},(k)}\right)+\mathbf{D}\left(\boldsymbol{\psi}_{l}^{n_{,}(k)}\right)-\mathbf{U}\left(\boldsymbol{\psi}_{l+1}^{n_{1}(k)}\right)-\rho_{l} \mathbf{R}_{l}^{n}\right\|_{\infty^{\prime}}, \tag{3.23}
\end{equation*}
$$

and choose to stop Algorithm 1 as soon as $k$ satisfies

$$
\begin{equation*}
r^{n,(k)} \leq r_{\max } \tag{3.24}
\end{equation*}
$$

Fig. 2 depicts the minimum number of iterations $k$ required in Algorithm 1 such that the the residual $r^{n,(k)}$ satisfies the criterium (3.24) with $r_{\text {max }}$ fixed at $1,10^{-1}$ and $10^{-2}$. This is represented as a function of the energy step $n$.


Figure 1: Normalized dose obtained with Algorithm 1 for the beam of electron (3.21).


Figure 2: Number of iterations $k$ as a function of the energy step $n$ for a given maximum residual $r_{\text {max }}$.


Figure 3: Final residual $r^{n,\left(k_{\max }\right)}$ as a function of the energy step $n$ for a given number of iteration $k_{\text {max }}$.

Similarily, Fig. 3 depicts the final residual $r^{n,\left(k_{\max }\right)}$ obtained by fixing $k_{\max }$ to 30,50 or 70 as a function of the energy step $n$.

In the first steps where the values of $\psi_{e}^{b}$, Algorithm 1 requires a large number of iterations to converge. At those energy, $\partial_{\epsilon} \psi \equiv\left(\psi^{n-1}-\psi^{n}\right) / \Delta \epsilon^{n}$ is large. Thus the intialization $\psi^{n,(0)}=\psi^{n-1}$ of Algorithm 1 is far from the desired solution. Therefore the present algorithm requires more iterations to converge.

The convergence rate progressively raises, i.e. the final residual $r^{n,\left(k_{\max }\right)}$ or the number of iterations $k$ reduce.

The drop of $r^{n,\left(k_{\max }\right)}$ and of iterations $k$ near the end of the simulation is due to the physical parameters used. The stopping power $S$ skyrockets near the threshold $\epsilon=\epsilon_{\min }$. According to the definition of $\mathbf{D}$ in (3.13), this implies that the eigenvalues of $\mathbf{D}$ also raise
at the lowest energy. Thus the eigenvalues of $J$ drop to zero, and so the convergence rate of Algorithm 1 skyrockets. This explains the shape of the curves $k$ and $r^{n,\left(k_{\max }\right)}$ in the last steps $n$.

In all the remaining test cases, the parameter $r_{\text {max }}$ is fixed, and the parameter $k_{\text {max }}$ is chosen sufficiently high such that the residual $r_{\max }$ is always reached during the computations, and only converged results are shown.

### 3.5.2 Convergence results of the numerical scheme

In this subsection, we verify experimentally that the numerical scheme is converging. The maximum residual $r_{\text {max }}$ was fixed at $5.10^{-3}$.

Since no analytical solution is known, the reference solution is chosen to be the solution obtained with the largest number of cells. The spatial domain $Z=[0,8 \mathrm{~cm}]$ is uniformly meshed. The number of spatial cells $l_{\max }$ is chosen to be 100, 200, 400, 800 and 1600 cells and for the reference solution 3200 cells. The convergence rate in $\Delta x$ is represented by the discrete $L^{1}, L^{2}$ and $L^{\infty}$ errors between the reference solution, i.e. the most refined one, and the less refined ones

$$
\begin{gathered}
\operatorname{Error}_{L^{1}}(\Delta x)=\sum_{l=1}^{l_{\max } n_{\max }} \sum_{n=1}^{n}\left|\psi_{l}^{n}-\tilde{\psi}_{l}^{n}\right| \tilde{\Delta x \Delta \tilde{\epsilon}^{n}}, \quad \operatorname{Error}_{L^{2}}(\Delta x)=\sqrt{\sum_{l=1}^{l_{\text {max }} n_{m a x}} \sum_{n=1}\left(\psi_{l}^{n}-\tilde{\psi}_{l}^{n}\right)^{2} \tilde{\Delta x} x \tilde{\epsilon}^{n}}, \\
\operatorname{Error}_{L^{\infty}}(\Delta x)=\max _{l=1, \ldots, l_{\max } n=1, \ldots, n_{\max }} \max _{l}\left|\psi_{l}^{n}-\tilde{\boldsymbol{\psi}}_{l}^{n}\right|,
\end{gathered}
$$

where $\psi_{l}^{n}$ is the solution with mesh cells of size $\Delta x$ and an energy step size $\Delta \epsilon^{n}$ given by (3.22) approximated by piecewise linear polynomials at the points ( $x_{l}, \epsilon^{n}$ ) on the finest mesh (of gridsize $\Delta x$ ) and $\tilde{\psi}_{l}^{n}$ is the most refined solution at the same points $\left(x_{l}, \epsilon^{n}\right)$. Those errors are plotted on Fig. 4 as a function of $\Delta x$. As expected, we observe a convergence


Figure 4: Discrete $L^{1}, L^{2}$ and $L^{\infty}$ errors compared to the most refined solution as a function of $\Delta x$ with $\Delta \epsilon^{n}$ given by (3.22).
rate of order 1, the mean slope of the curves obtained on Fig. 4 being of 1.230901 for the $L^{1}$ error, 1.213013 for the $L^{2}$ error and 0.974655 for the $L^{\infty}$ error.

## 4 A numerical approach for multi-D problems

We extend here the previous approach for coupled electrons and photons transport in multi-D media. The difficulties and the method are described in the next three subsection in 2D through the model

$$
\begin{equation*}
\partial_{x} \mathbf{F}_{1}(\boldsymbol{\psi})+\partial_{y} \mathbf{F}_{2}(\boldsymbol{\psi})=\rho \mathbf{Q}(\boldsymbol{\psi}) . \tag{4.1}
\end{equation*}
$$

However the method is also valid in 3D and a 3D test case is provided in Subsection 4.4.

### 4.1 A correction of the numerical transverse diffusion

As described in $[4,53]$ and through the experimental results below, using the relaxation parameters (3.9) when considering 2D photon beams leads to a numerical overestimation of the diffusion effects in the direction orthogonal to the beams.

In multi-D, the velocities $c^{ \pm}$are vectors instead of scalars. We consider two relaxation velocities $c_{i}^{ \pm}= \pm\left|c_{i}^{ \pm}\right| e_{i}$ per Cartesian direction $e_{i}$. The relaxation parameters need to satisfy ([1,11,45])

$$
\begin{align*}
& \forall d \in \mathcal{S}^{2}, \quad S p\left(\mathbf{F}_{d}^{\prime}(\boldsymbol{\psi})\right) \subset\left[\min _{i, \pm}\left(c_{i}^{ \pm} \cdot d\right), \max _{i, \pm}\left(c_{i}^{ \pm} \cdot d\right)\right],  \tag{4.2a}\\
& \sum_{i, \pm} \mathbf{M}_{i}^{ \pm}=\boldsymbol{\psi}, \quad \sum_{i, \pm}\left(c_{i}^{ \pm} \cdot d\right) \mathbf{M}_{i}^{ \pm}=\mathbf{F}_{d}(\boldsymbol{\psi}), \tag{4.2b}
\end{align*}
$$

where

$$
\mathbf{F}_{d}(\boldsymbol{\psi})=d_{1} \mathbf{F}_{1}(\boldsymbol{\psi})+d_{2} \mathbf{F}_{2}(\boldsymbol{\psi})=\left(\psi_{\gamma}^{1} \cdot d, \psi_{\gamma}^{2} \cdot d, \psi_{e}^{1} \cdot d, \psi_{e}^{2} \cdot d\right)
$$

is the flux in the direction $d$. In practice, we use the relaxation velocities $c_{i}^{ \pm}$defined in [53] that are assumed to approximate numerically the maximum physical velocities in each Cartesian directions

$$
\left|c_{i}^{ \pm}\right| \approx \max \left(\delta, \max \left[\operatorname{Sp}\left( \pm \mathbf{F}_{i}^{\prime}(\boldsymbol{\psi})\right)\right]\right)
$$

Here, $\delta=10^{-8}$ is a constant chosen arbitrarily small to avoid numerical divisions by zero. We use Maxwellians of the form

$$
\mathbf{M}_{i}^{ \pm}=\mu_{i}^{ \pm} \boldsymbol{\psi}+\lambda_{i}^{ \pm} \mathbf{F}_{i}(\boldsymbol{\psi}) .
$$

In order to satisfy (4.2b), one finds that $\mu_{i}^{ \pm}$and $\lambda_{i}^{ \pm}$need to satisfy

$$
\begin{gather*}
\sum_{i, \pm} \mu_{i}^{ \pm}=1,  \tag{4.3a}\\
\lambda_{i}^{-}+\lambda_{i}^{+}=0, \quad-\left|c_{i}^{-}\right| \mu_{i}^{-}+\left|c_{i}^{+}\right| \mu_{i}^{+}=0, \quad-\left|c_{i}^{-}\right| \lambda_{i}^{-}+\left|c_{i}^{+}\right| \lambda_{i}^{+}=1 . \tag{4.3b}
\end{gather*}
$$

The last three equations (4.3b) can be rewritten

$$
\mu_{i}^{-}=\frac{\left|c_{i}^{+}\right|}{\left|c_{i}^{-}\right|} \mu_{i}^{+}, \quad \lambda_{i}^{ \pm}= \pm \frac{1}{\left|c_{i}^{+}\right|+\left|c_{i}^{-}\right|} .
$$

In practice, in 2D, we choose to fix the last degrees of freedom by

$$
\mu_{i}^{ \pm}=\frac{\left|c_{i}^{\mp}\right|}{2\left(\left|c_{i}^{+}\right|+\left|c_{i}^{-}\right|\right)},
$$

which satisfies (4.3a). This leads to write the Maxwellians

$$
\begin{equation*}
\mathbf{M}_{i}^{ \pm}=\frac{\left|\tilde{c}_{i}^{\mp}\right| \boldsymbol{\psi} \pm 2 \mathbf{F}_{i}(\boldsymbol{\psi})}{2\left(\left|\tilde{c}_{i}^{+}\right|+\left|\tilde{c}_{i}^{-}\right|\right)} \in \mathcal{R}^{2}, \tag{4.4}
\end{equation*}
$$

where $\left|\tilde{c}_{i}^{ \pm}\right|$is either $\left|c_{i}^{ \pm}\right|$or the minimum scalar such that (4.4) is realizable.
Applying the method described in Subsection 3.2 to the 2D equation (4.1) leads to write the scheme

$$
\begin{gather*}
\frac{\mathbf{F}_{l+\frac{1}{2}, m}^{n}-\mathbf{F}_{l-\frac{1}{2}, m}^{n}}{\Delta x}+\frac{\mathbf{F}_{l, m+\frac{1}{2}}^{n}-\mathbf{F}_{l, m-\frac{1}{2}}^{n}}{\Delta y}-[\rho \mathbf{Q}(\boldsymbol{\psi})]_{l, m}^{n}=0,  \tag{4.5a}\\
\mathbf{F}_{l+\frac{1}{2}, m}^{n}=c_{l+\frac{1}{2}, m}^{-, n} \lambda_{l+\frac{1}{2}, m}^{-, n} \mathbf{F}_{1}\left(\boldsymbol{\psi}_{l+1, m}^{n}\right)+c_{l+\frac{1}{2}, m}^{+, n} \lambda_{l+\frac{1}{2}, m}^{+, n} \mathbf{F}_{1}\left(\boldsymbol{\psi}_{l, m}^{n}\right) \\
+\left(c_{l+\frac{1}{2}, m}^{-, n} \mu_{l+\frac{1}{2}, m}^{-, n} \boldsymbol{\psi}_{l+1, m}^{n}+c_{l+\frac{1}{2}, m}^{+, n} \mu_{l+\frac{1}{2}, m}^{+, n} \boldsymbol{\psi}_{l, m}^{n}\right),  \tag{4.5b}\\
\mathbf{F}_{l, m+\frac{1}{2}}^{n}=c_{l, m+\frac{1}{2}}^{-, n} \lambda_{l, m+\frac{1}{2}}^{-,} \mathbf{F}_{2}\left(\boldsymbol{\psi}_{l, m+1}^{n}\right)+c_{l, m+\frac{1}{2}}^{+, n} \lambda_{l, m+\frac{1}{2}}^{+, n} \mathbf{F}_{2}\left(\boldsymbol{\psi}_{l, m}^{n}\right) \\
 \tag{4.5c}\\
\quad+\left(c_{l, m+\frac{1}{2}}^{-, n} \mu_{l, m+\frac{1}{2}}^{-, n} \boldsymbol{\psi}_{l, m+1}^{n}+c_{l, m+\frac{1}{2}}^{+, n} \mu_{l, m+\frac{1}{2}}^{+, n} \boldsymbol{\psi}_{l, m}^{n}\right),
\end{gather*}
$$

where the coefficients in the discrete fluxes are given by

$$
\begin{gathered}
\lambda_{j, k}^{ \pm n}= \pm \frac{1}{\left|c_{j, k}^{+, n}\right|+\left|c_{j, k}^{-, n}\right|}, \quad \mu_{j, k}^{ \pm, n}=\frac{\left|c_{j, k}^{\mp, n}\right|}{2\left(\left|c_{j, k}^{+, n}\right|+\left|c_{j, k}^{-, n}\right|\right)}, \\
c_{l+\frac{1}{2}, m}^{ \pm, n}= \pm \max \left(\left|\tilde{c}_{1}^{ \pm}\left(\boldsymbol{\psi}_{l+1, m}^{n}\right)\right|,\left|\tilde{c}_{1}^{ \pm}\left(\boldsymbol{\psi}_{l, m}^{n}\right)\right|\right), \\
c_{l, m+\frac{1}{2}}^{ \pm, n}= \pm \max \left(\left|\tilde{c}_{2}^{ \pm}\left(\boldsymbol{\psi}_{l, m+1}^{n}\right)\right|,\left|\tilde{c}_{2}^{ \pm}\left(\boldsymbol{\psi}_{l, m}^{n}\right)\right|\right),
\end{gathered}
$$

for $(j, k)=\left(l+\frac{1}{2}, m\right)$ or $(j, k)=\left(l, m+\frac{1}{2}\right)$.

### 4.2 An iterative solver for the multi-D scheme with the transverse diffusion correction

We propose to adapt Algorithm 1 when the coefficients $c_{i}^{ \pm}$are not constants. In this case, the scheme (4.5) can be rewritten under the form

$$
\begin{align*}
\rho_{l} \mathbf{R}_{l}^{n}= & -\mathbf{L}_{1}\left(\boldsymbol{\psi}_{l-1, m}^{n}\right)-\mathbf{L}_{2}\left(\boldsymbol{\psi}_{l, m-1}^{n}\right)+\mathbf{D}\left(\boldsymbol{\psi}_{l, m}^{n}\right) \\
& -\mathbf{U}_{1}\left(\boldsymbol{\psi}_{l+1, m}^{n}\right)-\mathbf{U}_{2}\left(\boldsymbol{\psi}_{l, m+1}^{n}\right), \tag{4.6a}
\end{align*}
$$

where the operators $\mathbf{L}_{1}, \mathbf{L}_{2}, \mathbf{U}_{1}, \mathbf{U}_{2}$ and $\mathbf{D}$ yield

$$
\begin{align*}
\mathbf{L}_{1}\left(\boldsymbol{\psi}_{l-1, m}^{n}\right)= & \frac{c_{l-\frac{1}{2}, m}^{+, n}}{\Delta x}\left[\mu_{l-\frac{1}{2}, m}^{+, n} \boldsymbol{\psi}_{l-1, m}^{n}+\lambda_{l-\frac{1}{2}, m}^{+, n} \mathbf{F}_{1}\left(\boldsymbol{\psi}_{l-1, m}^{n}\right)\right],  \tag{4.6b}\\
\mathbf{L}_{2}\left(\boldsymbol{\psi}_{l, m-1}^{n}\right)= & \frac{c_{l, n-\frac{1}{2}}^{+, n}}{\Delta y}\left[\mu_{l, m-\frac{1}{2}}^{+, n} \boldsymbol{\psi}_{l, m-1}^{n}+\lambda_{l, m-\frac{1}{2}}^{+, n} \mathbf{F}_{2}\left(\boldsymbol{\psi}_{l, m-1}^{n}\right)\right],  \tag{4.6c}\\
\mathbf{U}_{1}\left(\boldsymbol{\psi}_{l+1, m}^{n}\right)= & \frac{-c_{l+\frac{1}{2}, m}^{-, n}}{\Delta x}\left[\mu_{l+\frac{1}{2}, m}^{-, n} \boldsymbol{\psi}_{l+1, m}^{n}+\lambda_{l+\frac{1}{2}, m}^{-, n} \mathbf{F}_{1}\left(\boldsymbol{\psi}_{l+1, m}^{n}\right)\right],  \tag{4.6d}\\
\mathbf{U}_{2}\left(\boldsymbol{\psi}_{l, m+1}^{n}\right)= & -c_{l, m+\frac{1}{2}}^{-n}\left[\mu_{l, m+\frac{1}{2}}^{-, n} \boldsymbol{\psi}_{l, m+1}^{n}+\lambda_{l, m+\frac{1}{2}}^{-, n} \mathbf{F}_{2}\left(\boldsymbol{\psi}_{l, m+1}^{n}\right)\right],  \tag{4.6e}\\
\mathbf{D}\left(\boldsymbol{\psi}_{l, m}^{n}\right)= & \left(\rho_{l} A^{n}+\beta_{l, m}^{n} I d\right) \boldsymbol{\psi}_{l, m}^{n}+\gamma_{1, l, m}^{n} \mathbf{F}_{1}\left(\boldsymbol{\psi}_{l, m}^{n}\right)+\gamma_{2, l, m}^{n} \mathbf{F}_{2}\left(\boldsymbol{\psi}_{l, m}^{n}\right), \tag{4.6f}
\end{align*}
$$

where the coeffiecients $\beta_{l, m}^{n}, \gamma_{1, l, m}^{n}$ and $\gamma_{2, l, m}^{n}$ read

$$
\begin{aligned}
\beta_{l, m}^{n} & =\frac{c_{l+\frac{1}{2}, m}^{+, n} \mu_{l+\frac{1}{2}, m}^{+, n}-c_{l-\frac{1}{2}, m}^{-, m} \mu_{l-\frac{1}{2}, m}^{-, n}}{\Delta x}+\frac{c_{l, m+\frac{1}{2}}^{+, n} \mu_{l, m+\frac{1}{2}}^{+, n}-c_{l, m-\frac{1}{2}}^{-, n} \mu_{l, m-\frac{1}{2}}^{-, n}}{\Delta y}, \\
\gamma_{1, l, m}^{n} & =\frac{c_{l+\frac{1}{2}, m}^{+, n} \lambda_{l+\frac{1}{2}, m}^{+, n}-c_{l-\frac{1}{2}, m}^{-, n} \lambda_{l-\frac{1}{2}, m}^{-, n}}{\Delta x}, \\
\gamma_{2, l, m}^{n} & =\frac{c_{l, m+\frac{1}{2}}^{+, n} \lambda_{l, m+\frac{1}{2}}^{+, n}-c_{l, m-\frac{1}{2}}^{-, n} \lambda_{l, m-\frac{1}{2}}^{-, n}}{\Delta y} .
\end{aligned}
$$

The difficulty here emerges from the non-linearity of the operator $\mathbf{D}$ to invert and from the realizability requirements (4.2).

Let us decompose the operator $\mathbf{D}$ into

$$
\begin{align*}
\mathbf{D}(\boldsymbol{\psi}) & =\mathbf{D}_{\text {imp }}(\boldsymbol{\psi})-\mathbf{D}_{\exp }(\boldsymbol{\psi}),  \tag{4.7}\\
\mathbf{D}_{\text {imp }}(\boldsymbol{\psi}) & =\left[\rho_{l} A^{n}+\left(\alpha_{l, m}^{n}+\beta_{l, m}^{n}\right)[d] \boldsymbol{\psi},\right. \\
\mathbf{D}_{\text {exp }}(\boldsymbol{\psi}) & =\alpha_{l, m}^{n} \boldsymbol{\psi}+\gamma_{1, l, m}^{n} \mathbf{F}_{1}(\boldsymbol{\psi})+\gamma_{2, l, m}^{n} \mathbf{F}_{2}(\boldsymbol{\psi}),
\end{align*}
$$

such that $\mathbf{D}_{i m p}$ is linear and invertible. Here the coefficient $\alpha_{l, m}^{n}$ is chosen non-negative such that the operator $\mathbf{D}_{\text {exp }}$ preserves the realizability. In practice, one may simply chose

$$
\alpha_{l, m}^{n}=\left|\gamma_{1, l, m}^{n}\right|+\left|\gamma_{2, l, m}^{n}\right| .
$$

Finally, Algorithm 1 is rewritten by modifying (3.14). This leads to the following algorithm.
Algorithm 2. Initialization: Set $\psi_{l, m}^{n,(0)}=\psi_{l, m}^{n-1}$ for all $l, m$. Iteration: Compute iteratively

$$
\begin{align*}
\boldsymbol{\psi}_{l, m}^{n+1,(k+1)}=\mathbf{D}_{i m p}^{-1} & \left(\mathbf{R}_{l, m}^{n}+\mathbf{L}_{1}\left(\boldsymbol{\psi}_{l-1, m}^{n+1,(k)}\right)+\mathbf{L}_{2}\left(\boldsymbol{\psi}_{l, m-1}^{n+1,(k)}\right)\right.  \tag{4.8}\\
& \left.+\mathbf{D}_{\exp }\left(\boldsymbol{\psi}_{l, m}^{n+1,(k)}\right)+\mathbf{R}_{1}\left(\boldsymbol{\psi}_{l-1, m}^{n+1,(k)}\right)+\mathbf{R}_{2}\left(\boldsymbol{\psi}_{l, m-1}^{n+1,(k)}\right)\right),
\end{align*}
$$

until convergence.
Remark 4.1. - The parameter $\alpha^{n} I d$ was artificially added on both sides of (4.7) when splitting the operator $\mathbf{D}$ in two parts. This enforces the preservation of the realizability property and makes the algorithm more stable. However, this reduces the convergence rate of Algorithm 2.

- One may reproduce the computations from the proof of Proposition 3.1 with (4.8) to show that the realizability property is preserved from one iteration to another in Algorithm 2 and to show it is convergent.


### 4.3 Numerical experiment in 2D: a photon beam in water

In this test case, photons are injected in a 2D homogeneous domain composed of water. The size of the medium is $2 \mathrm{~cm} \times 10 \mathrm{~cm}$, and a 0.5 cm large beam of 500 keV photons is injected on the left boundary. This is modeled by the following incoming boundary condition

$$
\text { for } \begin{aligned}
(X, \Omega) \in \Gamma^{-} & =\left\{(X, \Omega) \in \partial Z \times \mathcal{S}^{2} \quad \text { s.t. } \quad n(X) . \Omega<0\right\}, \\
\psi_{\gamma}(X, \epsilon, \Omega) & =10^{10} \exp \left(-\alpha_{\epsilon}\left(\epsilon-\epsilon_{0}\right)^{2}\right) \exp \left(-\alpha_{\mu}\left(\Omega_{1}-1\right)^{2}\right) \mathbf{1}_{B}(X)+\delta \mathbf{1}_{\partial Z \backslash B}(X), \\
\psi_{e}(X, \epsilon, \Omega) & =\delta, \\
B & =\{(x, y), \quad x=0, \quad y \in[0.75 \mathrm{~cm}, 1.25 \mathrm{~cm}]\},
\end{aligned}
$$

where $n(X)$ is the outgoing normal, $\epsilon_{0}=500 \mathrm{keV}, \alpha_{\epsilon}=20000, \alpha_{\mu}=500000$ and $\delta=10^{-15}$. And we used the moments of those distribution as boundary conditions for the moment equations.

Through this test case, we aim to highlight the influence of the choice of the parameters $c^{ \pm}$on the convergence rate of Algorithm 1. The influence of $c^{ \pm}$on the dose results were studied in [53] and are only recalled here for completeness. The present method is tested using two sets of parameters $c_{i}^{ \pm}$:

- First, we fix $\left|c_{i}^{ \pm}\right|=2$ at a sufficiently large value (afterward called large $c$ ) for the conditions (4.2) to be satisfied. Those parameters were shown to provide an overestimated numerical angular diffusion in [53] (see also Fig. 5 below).
- Second, we fix $c_{i}^{ \pm}$at a value closer to the value actual eigenvalues of the Jacobian of the flux as proposed in [53] (afterward called small c). These parameters were shown to reduce the diffusion effects of the numerical method, especially in the direction orthogonal to the beam.

Algorithm 2 is compared to a reference Monte Carlo solver ( [24]). The dose results obtained with those methods are gathered on Fig. 5 with the computational times in Table


Figure 5: Doses obtained with the Monte Carlo solver (top) and the $M_{1}$ solver with large $c$ (middle) and small $c$ (below) relaxation parameters, normalized by their maximum value.

| Solver | Monte Carlo | $M_{1}$ with large $c$ | $M_{1}$ with small $c$ |
| :---: | :---: | :---: | :---: |
| Computation times | 14 hours | 49.78699 sec | 204.1239 sec |

Table 1: Computational times with the Monte Carlo solver and the implicit solver with the different $c$.

1. A cut of the doses along the axis of the beam $y=1 \mathrm{~cm}$ and in the transverse direction at depth $x=2 \mathrm{~cm}$ and $x=8 \mathrm{~cm}$ are shown on Fig. 6 .

The dose results with the modified relaxation parameters are much closer to the reference Monte Carlo results. As expected the dose is less diffused with the small $c$.

Due to the noise in the Monte Carlo and the normalization by $\max D$, the $M_{1}$ dose curves with the modified relaxation parameters are slightly above the Monte Carlo reference on Fig. 6.

Following Remark 4.1, we observe that the computational time is higher with the small $c$ than with the large ones. Those times remains much lower than the one with the Monte Carlo reference.


Figure 6: Doses obtained with the Monte Carlo solver and the $M_{1}$ solver with large and small $c$ along the axis of the beam (top) and the axis transverse to the beam at 2 cm depth (below left) and 8 cm depth (below right), normalized by their maximum value.

### 4.4 Numerical experiment in 3D: a photon beam in a chest

This test case aims to exhibit the efficiency of our method when considering more complex density maps. For this purpose, a density map obtained from a computed tomography (CT) scan of a chest is used. This map is depicted on Fig. 7. This domain is a 29.5 cm deep cube.

A beam of photons is imposed on the boundary of the medium to pass through the ribs. In order to reduce the computational time, the computations are performed on a smaller domain of size $14 \mathrm{~cm} \times 25 \mathrm{~cm} \times 11.35 \mathrm{~cm}$ in which the density of photons is non-negligible. This domain is meshed with $140 \times 220 \times 50$ cells.


Figure 7: Density map represented by isosurfaces of density 1.8 (ivory; equivals to bone density) and 0.3 (flesh colour; equivals to lung density).

The beam is modeled by the following condition over $\Gamma^{-}$

$$
\begin{aligned}
\psi_{\gamma}(X, \epsilon, \Omega) & =10^{10} \exp \left(-\alpha_{\epsilon}\left(\epsilon-\epsilon_{0}\right)^{2}\right) \exp \left(-\alpha_{\mu}\left(\Omega_{2}-1\right)^{2}\right) \mathbf{1}_{B}(X)+\delta \mathbf{1}_{\partial Z \backslash B}(X), \\
\psi_{e}(X, \epsilon, \Omega) & =\delta, \\
B & =\{(x, y, z), \quad x \in[6 \mathrm{~cm}, 8 \mathrm{~cm}], y=0 \mathrm{~cm}, z \in[4 \mathrm{~cm}, 6 \mathrm{~cm}]\}
\end{aligned}
$$

where $n(x)$ is the outgoing normal, $\epsilon_{0}=1 \mathrm{MeV}, \alpha_{\epsilon}=20000, \alpha_{\mu}=3000$ and $\delta=10^{-15}$. And we used the moments of those distribution as boundary conditions for the moment equations.

The maximum residual was fixed at $r_{\max }=10^{-1}$.
The computations were performed with the small parameters $c$. Those dose results are depicted on Fig. 8 as isodose surfaces cut along the surfaces parametrized by $\{x=5$ $\mathrm{cm}\},\{y=12.5 \mathrm{~cm}\}$ and $\{z=5.675 \mathrm{~cm}\}$, that are along the axis of the beam or at half depth in the domain.

The present method was tested on a rather complex problem with a large 3D mesh, a complete physics (photons and electrons together) and using the large $c$, that is in the most complex settings. The computations were performed in parallel on four cores and the computational times are gathered in Table 2. These times of computataion remains


Figure 8: Isosurfaces of dose at $60 \%, 30 \%, 6 \%, 3 \%$ and $0.6 \%$ of the maximum dose in a chest.

| Solver | $M_{1}$ with small $c$ |
| :---: | :---: |
| Computation times | $10405 \mathrm{sec} \approx 2 \mathrm{~h} 53 \mathrm{~min}$ |

Table 2: Computational times with the implicit solver for the 3D problem.
too long for practical applications in medical physics. Several features in our approach can be improved to reduce drastically the computational times. The main time consuming issue and an axe of improvement are presented here.

In practice, the moments $\psi$ are computed in the whole space-energy domain, even in cells where they are negligible. This results in solving (4.8) over a very large vector $\psi^{n}$. Those non-necessary computations raise the size of the system to solve and reduce the convergence rate of our algorithm (see Remark 3.3).

The computational times can be considerably reduced using common code optimization techniques. The most common idea for such a problem is the mesh adaptation, here over the space and the energy. Instead of computing $\psi$ in the whole domain, one may solve (4.8) on a smaller domain where $\psi$ is expected not to be negligible. This would lead to a considerable reduction of both the computational times and the memory requirements. This method was also used in the development of the industrial code Acuros ${ }^{\circledR}[25,46,60]$ with considerable ressources requirement reduction.

## 5 Conclusions

The present approach aims to circumvent restrictive stability conditions when numerically solving coupled linear kinetic equations.

First, an angular moment extraction was performed leading to a so-called $M_{1}$ system of equations. Such models require lower computational cost to solve, but are valid under a realizability condition.

A numerical scheme for $M_{1}$ systems was proposed by chosing implicit flux and collision terms. Then an iterative method was constructed to compute the solution of such an implicit scheme. The proposed algorithm was constructed with a special focus on the preservation of the realizability property.

Numerical experiments showed that our method behaves appropriately in practical cases in 1D and in multi-D and the convegence of the method was tested in 1D. The present approach was shown to require much lower computational costs than a reference Monte Carlo method.

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